

# IP2: Image Processing in Remote Sensing

# 3. Electromagnetic Radiation I: Waves and Basic Principles

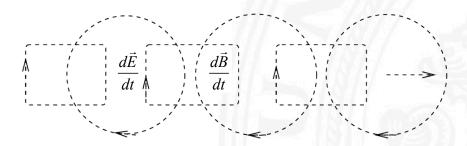
Summer Semester 2014 Benjamin Seppke

# **Agenda**

- Electromagnetic Waves
- Harmonic Oscillation
- Interaction with Transmission Medium
- Wave packages
- The Doppler Effect
- Other Relativistic Effects

## **Electromagnetic Radiation**

- EM Radiation is the most important transmission medium for Remote Sensing applications
- Consists of electric and magnetic fields (Maxwell's Equations)
- Creation of spatial EM waves:
  - A temporal varying electric field creates a magnetic field
  - A temporal varying magnetic field creates an electric field
    - → A traversal wave is created.



*E*: electric field strength

elektrische Feldstärke

**B**: magnetic field intensity

magnetische Feldstärke

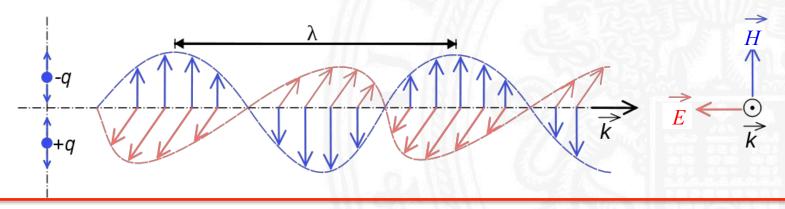
## **EM Spatial Waves**

#### A planar EM wave consist of:

- An electric field, which oscillates perpendicular to the travelling direction k. It is described by the vector of electric field strength E.
- A magnetic field, perpendicular to the electric field and oscillating perpendicular to the travelling direction k. It is described by the vector of magnetic field intensity H.

#### Characterization of the EM wave:

- Wavelength:  $\lambda$
- Frequency: ν
- Velocity:  $v = \lambda v$  (at vacuum: c)



## **Maxwell's Equations**

#### Following Rees, 1990:

$$\nabla \cdot \vec{E} = 0$$

$$\nabla \cdot \vec{\mathbf{B}} = 0$$

$$\nabla \times \vec{E} = -\frac{\partial \vec{B}}{\partial t}$$

$$\nabla \times \vec{\boldsymbol{B}} = \frac{\mu_r \varepsilon_r}{c^2} \frac{\partial \vec{\boldsymbol{E}}}{\partial t}$$

with:

Inductivity:  $\vec{B} = \mu_0 \mu_r \vec{H}$ 

and material properties:

Dielectric constant:  $\varepsilon$ 

 $\blacksquare$  Dielektrizitätskonstante: arepsilon

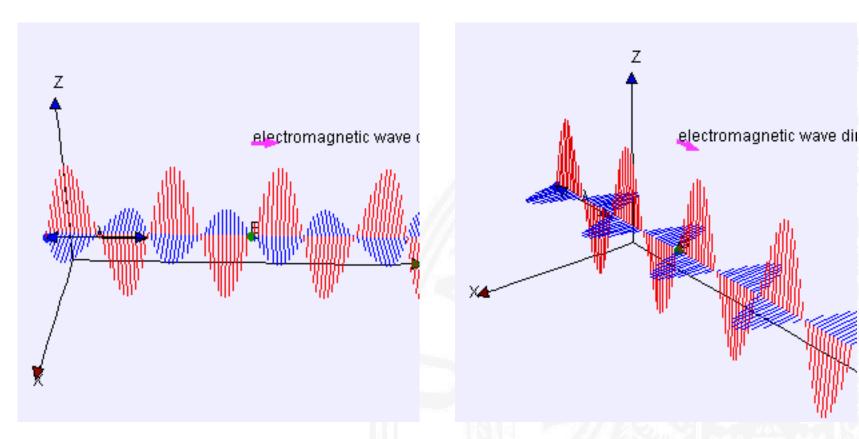
Permeability:  $\mu$ 

**—** Permeabilität: μ

#### Notation for material property coefficients:

- Inside vacuum:  $\varepsilon_{ heta}$  ,  $\mu_{ heta}$
- Inside a medium (relative to vacuum):  $\varepsilon_r$  ,  $\mu_r$  or:  $\varepsilon$  ,  $\mu$

## Maxwell's Equations: Live-Demo



#### **Created with:**

https://dl.dropboxusercontent.com/u/44365627/lookangEJSworkspace/export/ejs\_users\_sgeducation\_lookang\_emwavewee.jar

## **General Properties of EM Waves**

- Die planar spatial wave is a transversal wave.
- The electric field E is perpendicular to the magnetic field H(B) the travelling direction of the wave  $E \times H(E \times B)$  and is perpendicular to both fields
- The oscillation plane of the *H*-vector is called polarization plane of the wave.
- The oscillation plane of the *E*-vector is called oscillation plane of the wave..
- The vector  $E \times B$  is called Pointing vector S.
- The energy transport of the wave in performed in the direction of S:

 $|S| = E \cdot B$  gives the energy, which is transported perpendicular to the travelling direction. (in time units per area units).

# **Planar EM Waves Equation**

Given as a solution of Maxwell's Equations in homogeneous isotropic, non-magnetic media with the following wave equations:

$$\nabla^2 \vec{E} - \mu_0 \varepsilon_0 \mu_r \varepsilon_r \frac{\partial^2 \vec{E}}{\partial t^2} = 0$$

$$\nabla^2 \vec{B} - \mu_0 \varepsilon_0 \mu_r \varepsilon_r \frac{\partial^2 \vec{B}}{\partial t^2} = 0$$

with (partial) spatial derivatives:

$$\nabla^2 \vec{E}, \nabla^2 \vec{B}$$

and (partial) temporal derivatives:

$$\frac{\partial^2 \vec{\boldsymbol{E}}}{\partial t^2}, \frac{\partial^2 \vec{\boldsymbol{B}}}{\partial t^2}$$

**Note:** Temporal change of a wave at a location yields to a change of the wave in travelling direction.

# **Harmonic EM Waves (1)**

For harmonic waves, the wave equation is given by:

$$\nabla^{2}\vec{E} - \frac{\omega^{2}}{c_{r}^{2}}\vec{E} = 0$$
 with: 
$$v = c_{r} = \frac{c_{0}}{\sqrt{\mu_{0}\varepsilon_{0}\mu_{r}\varepsilon_{r}}}$$
 
$$c = c_{0} = \frac{1}{\sqrt{\mu_{r}\varepsilon_{r}}}$$

Using Euler's Formula, we can express the wave function complex-valued:

 $\vec{E} = \vec{A}e^{i(kr - \omega t + \phi)}$ 

and:

#### with:

- $\vec{A}$  Amplitude
- $\omega$  Complex angular frequency
- *t* Time point
- k Wave vector at medium
- <sup>r</sup> Space point
- $\phi$  Phase

#### Euler's Formula

$$\cos x = \frac{1}{2} \left( e^{ix} + e^{-ix} \right)$$

$$\sin x = \frac{1}{2i} \left( e^{ix} - e^{-ix} \right)$$

- $k = \frac{2\pi\sqrt{e_r}}{\lambda}$
- $\lambda = \frac{2\pi c_0}{\omega}$  Wavelength
- $v = \frac{\omega}{2\pi}$  Frequency

# Harmonic EM Waves (2)

**Another definition:** The wave can be described by a sine in a three dimensional space: The electric field strength E is parallel to the x-axis, the vector B is parallel to the y-axis, the travelling direction of the wave z.

with:

$$\vec{E}_{y} = \vec{E}_{z} = 0$$

$$\vec{E}_{x} = \vec{E}_{0} \cos(\omega t - kz)$$

$$\vec{B}_{x} = \vec{B}_{z} = 0$$

$$\vec{B}_{y} = \vec{B}_{0} \cos(\omega t - kz)$$

$$\omega = 2\pi v = \frac{ck}{n}$$
 Angular frequency  $k = \frac{2\pi}{\lambda}$  Wave number  $v = \frac{c}{\lambda n}$  Frequency  $v = \lambda v = \frac{c}{n}$  Phase velocity  $n = \sqrt{\varepsilon \mu}$  Refractivity Brechungsindex

- $E_0$ ,  $B_0$  are amplitudes of field intensities
- kz space dependent variation of field intensity
- $\omega t$  space dependent variation of field intensity

## The Transmission Medium

- General properties
- Impedance
  - Wellenwiderstand, Impedanz
- Dielectric constant
  - **—** Dielektrizitätskonstante
    - Absorbing media
    - Dispersing media
    - Plasma

## **Transmission Medium: General properties**

- $c_r$  (or v) denotes the phase velocity of a wave inside the medium,
- $c_0$  (or c or C) = 2.99792458×10<sup>8</sup> m/s the phase velocity of a wave inside vacuum.
- The permeability  $\mu_r$  (usually) equals 1 (for non-magnetic media)
- The dielectric constant  $\varepsilon_r$  (usually) varies between 1 to 80. [Elachi, 1987]
  - $\varepsilon_r$  is a function of the frequency  $\rightarrow$  dispersion!

## **Transmission Medium: Impedance**

- The electric and the magnetic field intensity are oscillating with same phase, thus the ratio of both field intensities (the impedance) remains constant.
- The impedance Z is a characteristic constant of a medium.
- $Z_0$  denotes the "impedance at vacuum".

$$Z = \frac{E}{H} = \frac{E}{B} \mu_r \mu_0 = \frac{E_0}{B_0} \mu_r \mu_0 = v \cdot \mu_r \mu_0$$
$$= \sqrt{\frac{\mu_r \mu_0}{\varepsilon_r \varepsilon_0}} = Z_0 \sqrt{\frac{\mu_r}{\varepsilon_r}}$$
$$Z_0 = \sqrt{\frac{\mu_0}{\varepsilon_0}} = 377\Omega$$

• Impedance is inverse proportional to the phase velocity of the wave: The higher the impedance, the slower the wave!

## **Transmission Medium: Dielectric constant**

• **Remember:** The phase velocity v of a spatial wave depends on the dielectric constant  $\varepsilon$  of the medium:

$$v = \frac{c}{n}, \quad n = \sqrt{\varepsilon \mu}$$

- Assuming  $\mu = 1$  it follows that  $n = \sqrt{\varepsilon}$
- Attention! Only valid for lossless transmission media the wave needs to pass the medium without loss of energy.

## **Dielectric Constant in Absorbing Media**

• In absorbing media, the dielectric constant  $\varepsilon$  is complex valued:

$$\varepsilon = \varepsilon' - i \cdot \varepsilon''$$

And so is the refractivity:

$$n = n' - i \cdot \kappa$$

• Alternate notation (with  $tan(\delta)$  - loss tangent):

$$n = \varepsilon' \Big( 1 - i \cdot \tan(\delta) \Big)$$

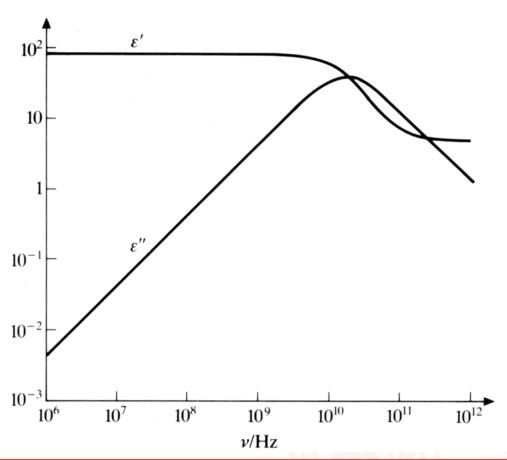
Verlusttangens

## Dielectric Constant in Dispersive Media

- In non-polarizing media,  $\varepsilon$  and  $\varepsilon$  are constant.
- If the medium is polarizing,  $\varepsilon$  and  $\varepsilon$  are (as well as the refractivity n) depending on the frequency of the wave.
  - → Dispersion, e.g. water, glass
- The frequency dependency of dispersive media is expressed by the Debye Equation (resonance effects, see Rees 1990)
- In electrically conducting medias, the imaginary component of the refraction index depends on the conductivity of the material.

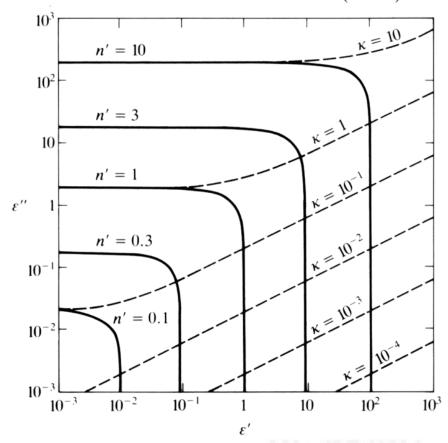
## **Example: Refraction Index of Water**

Fig. 2.2. The real and imaginary parts of the dielectric constant of pure water at 20 °C.



## Dielectric constant vs. refraction index

Fig. 2.3. The relationship between dielectric constant  $(\varepsilon' - i\varepsilon'')$  and refractive index  $(n' - i\kappa)$ .



## **Dielectric Constant in Plasma**

- In plasma (e.g. inside the Ionosphere) all atoms are ionized.
- Let:

N be the count of ions per volume unit m be the mass of the particles:

$$\varepsilon = n^2 = 1 - N \frac{e^2}{\varepsilon_0 m \omega^2}$$

#### **Observations:**

- *n* is real valued for high frequencies and imaginary valued fro low frequencies
- $\varepsilon$  may become smaller than 1.
  - $\rightarrow$  The phase velocity becomes larger than c (speed of light)!

## **Waves Packages**

Wellengemische, Wellenpakete

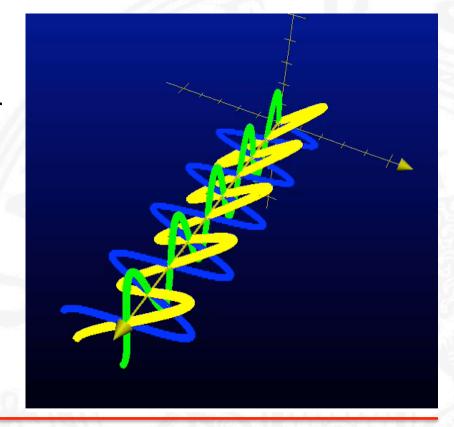
## From single waves $\rightarrow$ wave packages:

- Polarization
  - Linear
  - Circular
  - Elliptic
- Coherence
- Wave packages' velocities
  - Phase velocity
  - Group velocity

#### **Linear Polarization**

An EM wave is linear polarized, if the electric field E oscillated within a fixed plane.

- Horizontal Polarization:
   The vector *E* is perpendicular to the plane of incidence of the radiation.
- Vertical Polarization:
   The vector *E* lies within the plane of incidence of the radiation.

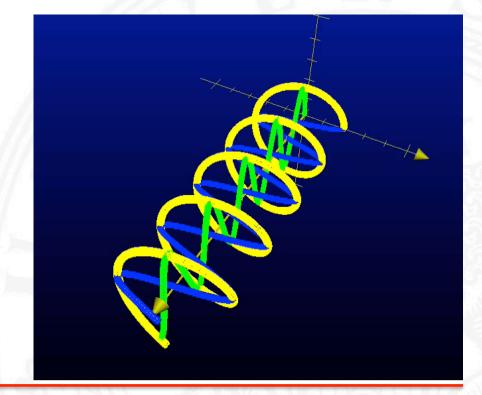


## **Circular Polarization**

The superposition of two polarized waves of **equal** phase and wavelength yields to a new **linear polarized** wave. The polarization plane is given by the addition of both electric

fields.

 The superposition of two linear polarized waves of different polarization and phase yields to a circular polarized wave, with a rotating direction of the electric field



## **Further Polarization Properties**

- The relative phase of both components determines if the resulting wave is:
  - Circular (shape of the helix is a circle)
  - Elliptic (shape of the helix is an ellipse) polarized
- Depending on the rotational direction is the wave called:
  - Left-hand (LHC) polarization (counterclockwise)
  - Right-hand (RHC) polarization (clockwise)
- Radiation is named non-polarized, if the polarization randomly varies
- Some sensors are only sensitive to certain polarization schemes, e.g. H- and V-Glasses for 3D movies at cinema

## **Coherence**

The coherence of a wave package is the time, which passes between the amplitudes of strong correlations. More precisely:

• For two waves with frequencies v and  $v + \Delta v$  the coherence time  $\Delta t$  is defined as the amount of time, after which the phases are shifted one cycle at each other:

$$v \Delta t + 1 = (v + \Delta v) \Delta t$$

$$\Rightarrow \Delta v \Delta t = 1$$

$$\Rightarrow \Delta t = \frac{1}{\Delta v}$$

The coherence length is defined as:

$$\Delta l = c \, \Delta t = \frac{c}{\Delta v}$$

## **Phase and Group Velocity**

• The phase velocity  $v_p$  of a wave is the velocity at which the phase of the wave propagates in space:

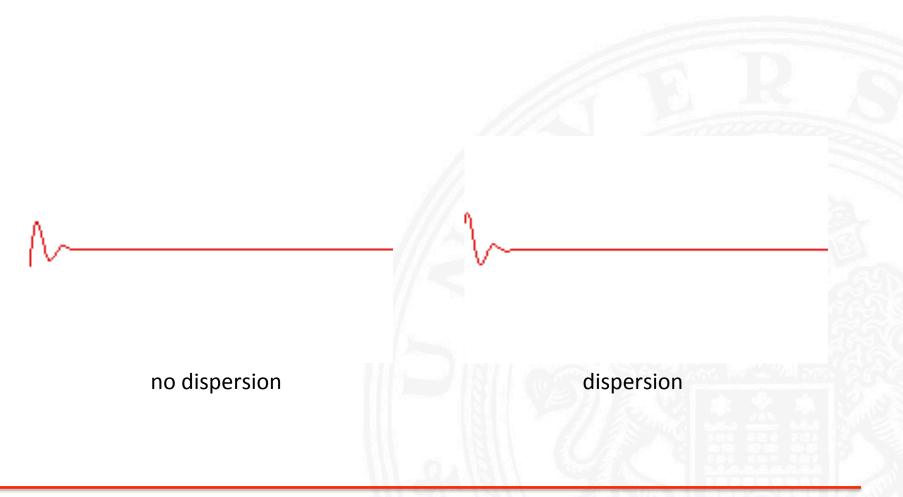
 $v_p = \frac{\omega}{k}$ 

• Sensors (often) do not measure the phase velocity but the group velocity  $v_g$ , which is the velocity of pulses, which are modulated on carrier waves:

 $v_g = \frac{\mathrm{d}\omega}{\mathrm{d}k}$ 

- The modulations results in additional frequencies. In dispersive media  $v_g < v_p$  holds.
- $v_g$  is the velocity, which is used to transport energy. Thus  $v_g \leq c$  holds, too.

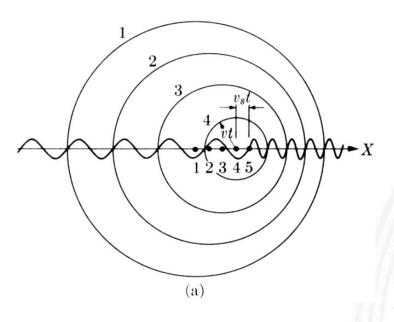
# Phase and Group Velocity (2)



## **Doppler Effect**

- The result of the observation of a wave's frequency depends on the velocity and direction of the observer, the sender and the transmission medium:
  - Moving sender
  - Moving observer
  - Moving reflecting, scattering body
- For large velocities (close to c), the Doppler Effect needs to be computed with the equations of the special relativity.
- For spatial EM waves, the Doppler effect depends only on the relative velocity between observer and sender
- For acoustic or surface waves other equations may be used for computation, depending on the motion of observer or sender or both.

## **Moving Sender**



- The apparent wavelength is shortening in motion direction.
- The apparent wavelength is increasing opposite to the motion direction.
- Reason: velocity of sender adds to the velocity of the wave.
- For an observer from the side, only the velocity component of the sender in direction to the observer is "visible".

## **Doppler Effect for EM Radiation**

#### High velocities require relativistic formulas:

$$\frac{v'}{v} = \frac{\sqrt{1 - \frac{V^2}{c^2} \cos^2(\Theta)}}{1 - \frac{V}{c} \cos(\Theta)}$$

with: *v* emitted frequency

*v* ' observed frequency

V relative velocity of sender S and observer O

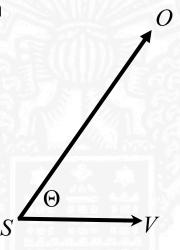
Θ Angle between motion and observation direction

#### (Very) small velocities: non-relativistic formula:

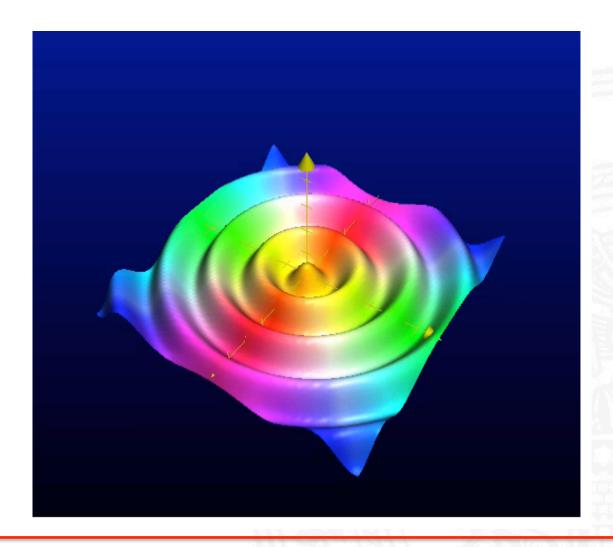
$$\frac{v'}{v} = 1 + \frac{V}{c}\cos(\Theta)$$

#### **Example:**

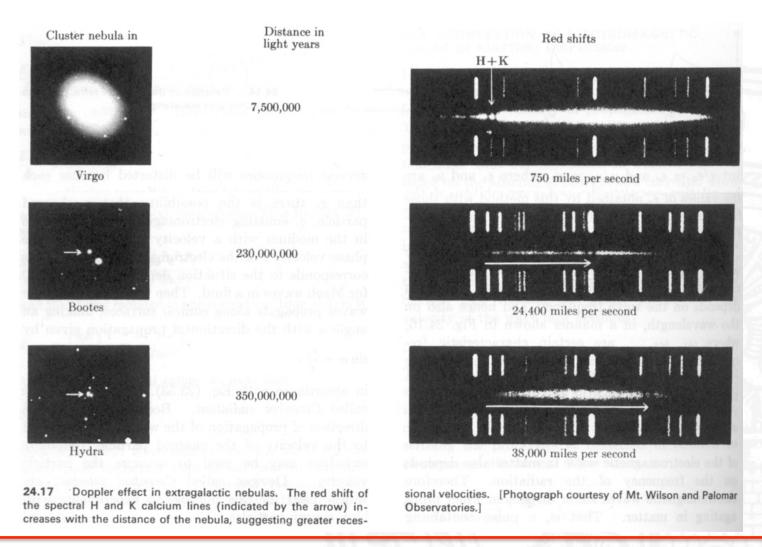
Satellite V = 7 km/s,  $\Theta = 5^{\circ}$ ,  $v = 5 \text{GHz} \rightarrow v - v' = 116 \text{ kHz}$ 



# Doppler Effect: 0.1c - 0.9c

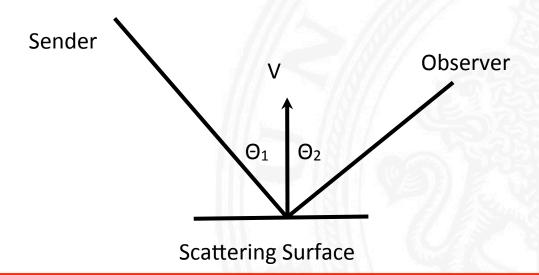


#### Relativistic Doppler Effect: The Expanding Universe



# **Moving Reflector and Scattering (1)**

- The Doppler Effect also appears, if neither the signal source nor the sensor is moving!
   Precondition: the wave is scattered by a third object or is directly reflected from a moving object!
- Typical situation in RADAR-based Remote Sensing



# **Moving Reflector and Scattering (2)**

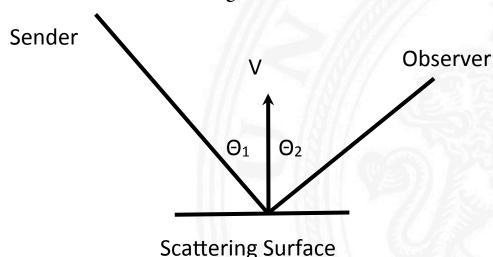
Computation of the increasing rate of the frequency vd

Observer and sender at different locations:

$$vd = v(\cos(\Theta_1) + \cos(\Theta_2))\frac{v}{c}$$

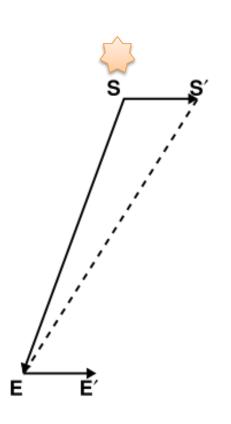
Observer and sender at the same location:

$$\Theta_1 = \Theta_2 = \Theta \Rightarrow vd = 2v\cos(\Theta_1)\frac{v}{c}$$



## **Aberration of Light**

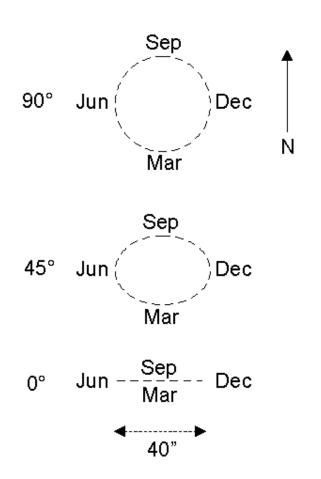
Abberation des Lichts

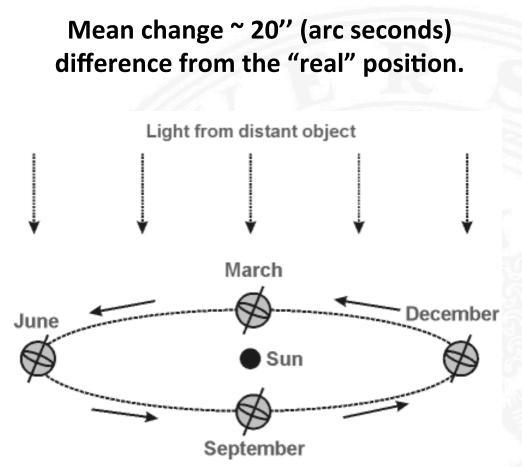


If observer and light source are moving relative to each other:

- not only the apparent frequency is changing but
- the apparent position of the light source is changing, too!

## The Seasonal Aberration





## Time of Flight

- For position measurements at large distances, the time of flight (of light) needs to be taken into account.
- We measure the position at the time of their emission, not at the arrival of the light.
- This can yield to large effects w.r.t. planetary space probes:
- Exemplary Runtimes:
  - Earth → Moon: ca. 1 second
  - Earth  $\rightarrow$  Sun: ca. 8.2 minutes